values of  $p_{\rm v}/p_{\rm b}$  are observed in Cygnus and Cepheus-Cassiopeia; for longitudes  $140^\circ < l^{\rm H} < 216^\circ$  the ratio is higher than that in Cygnus, although the largest value of  $p_{\rm v}/p_{\rm b}$  is found in Sagittarius.

There are two possible explanations to account for the variation of the extinction curves. At a given wavelength the cross-sections of interstellar elongated particles (aligned by a magnetic field) are different for different orientations; it has been predicted12,13 that the ratio

 ${{{\bf UV}~{
m slope}}\over {{\bf v}{{\bf v}}{{\bf v}}}}$  of the extinction curve would be dependent on BV slope

the angle of viewing and that this dependence would be related to polarization relative to extinction. Observation has not confirmed these predictions. The other possibility is that the composition and size distribution of interstellar dust grains could be different for different parts of the galaxy. If the relative number of large particles is increased, it is to be expected that a decrease in the ratio

of the extinction curve will be related to the BV slope

shift of the maximum polarization towards longer wavelengths and to an increase in the value of the ratio of total to selective absorption. Some confirmation of these predictions is afforded by our results and those of Serkowski.

McCuskey14 has found that the obscuration in the anticentre region is not very high and has concluded that there is a real deficiency of stars there. But he assumed a ratio of total to selective absorption  $R = A_{V}/E_{B-V}$  close to 3. A much higher value of R, or the addition of much grey obscuring matter in an anticentre cloud, would lead to underestimation of the total obscuration. This in turn would cause the distances of the stars to be overestimated and produce a spurious reduction in star density in the region. It may be significant that the volume density of stars derived by McCuskey drops suddenly at  $l^{II} = 140^{\circ}$ , the contours of equal surface density being almost radial to the Sun at this galactic longitude. This is true even for F0-F5 stars within 500 pc.

A high value of R, or grey obscuration, implies a cloud of particles with large sizes, and this may indicate that the grain temperature in the cloud is very low so that grains that were initially small have grown into large agglomera-Such a low temperature has interesting consequences in terms of the formation of molecules and the fragmentation and condensation of the cloud.

 $\check{A}$  discontinuity at  $l^{\rm H}=140^\circ$  has been found to occur sharply in the apparent distribution of OB stars 15 and to be inclined at about 45° to the Galactic Equator. A survey of a region 2° in radius centred on  $l^{II} = 140^{\circ}$ ,  $b^{II} = 0$ , and including UBVI photometry and polarization measures of all stars in the region brighter than about B=16, is being undertaken at this observatory.

M. T. Brück J. G. IRELAND K. Nandy V. C. REDDISH

Royal Observatory, Edinburgh.

Received April 16, 1968.

- Reddish, V. C., Nature, 213, 1107 (1967).
   Becker, W., Z. Für. Ap., 58, 202 (1964).
- Westerhout, G., Maryland-Greenbank Galactic 21-cm Line Survey, first ed. (University of Maryland, 1966).
- 4 Nandy, K., Pub. Roy. Obs. Edin., 3, 142 (1964).
- <sup>5</sup> Nandy, K., Pub. Roy. Obs. Edin., 5, 13 (1965).
- <sup>6</sup> Nandy, K., Pub. Roy. Obs. Edin., 5, 233 (1966).

- <sup>1</sup> Nandy, K., Pub. Roy. Obs. Edin., 6, 25 (1967).

  <sup>1</sup> Nandy, K., Pub. Roy. Obs. Edin., 6 (25 (1967).

  <sup>2</sup> Nandy, K., Pub. Roy. Obs. Edin., 6 (in the press).

  <sup>3</sup> Whitford, A. E., Astron. J., 63, 201 (1958).

  <sup>30</sup> Underhill, A. B., and Walker, G. A. H., Mon. Not. Roy. Astro. Soc., 131, 475 (1966).
- <sup>11</sup> Reddish, V. C., Pub. Roy. Obs. Edin., 6, 15 (1967).
- <sup>12</sup> Greenberg, J. M., and Meltzer, A. S., Astrophys. J., 132, 667 (1960).
- <sup>13</sup> Wilson, R., Mon. Not. Roy. Astro. Soc., 120, 51 (1960).
- McCuskey, S. W., Astrophys. J., 123, 458 (1956).
   Sim, M. E., Pub. Roy. Obs. Edin., 6 (in the press).

## BL Lac identified as a Radio Source

I wish to draw attention to the fact that a new position  $(\alpha_{1950} = 22h \ 00m \ 38s \cdot 9 \pm 0s \cdot 7, \ \delta_{1950} = + 42^{\circ} \ 02' \ 09'' \pm 9'')$  and optical identification of the radio source VRO 42.22.01 (ref. 2) determined with the Algonquin Radio Observatory 150 ft. telescope places it coincident with the irregular variable star BL Lac. BL Lac was discovered by Hoffmeister<sup>3</sup>, and varies from the thirteenth to sixteenth magnitude with fluctuations of large fractions of a magnitude in a few days4. A finding chart is published by Hoffmeister<sup>5</sup>; the position of the variable in the chart by Semakin<sup>4</sup> seems to be slightly in error. Examination of a glass copy of the National Geographic Society and Palomar Observatory Sky Survey at the Dunlap Observatory shows marginal nebulosity about the star. The optical properties of this object combined with its radio polarization and unusual microwave spectrum make it outstandingly interesting.

JOHN L. SCHMITT

David Dunlap Observatory, University of Toronto, Richmond Hill, Ontario, Canada.

Received April 24, 1968.

- MacLeod, J. M., and Andrew, B. H., Astrophys. Lett. (in the press).
   MacLeod, J. M., Swenson, jun., G. W., Yang, K. S., and Dickel, J. R., Astron. J., 70, 756 (1965).
- <sup>3</sup> Hoffmeister, C., Astron. Nachr., 236, 233 (1929). <sup>4</sup> Semakin, N. K., Variable Stars (Moscow), 10, 283 (1955).
- <sup>5</sup> Hoffmeister, C., Sonneberg Mitt., No. 17 (1930).

## Cosmological Significance of Time Reversal

A RECENT attempt by Stannard<sup>1</sup> to explain the apparent overthrow of parity in the long-lived kaon experiments suggested the possibility of an unseen component of the universe in which matter was of opposite time sense to that in the observable universe.

Beginning with this idea, a cosmological model with interesting symmetry properties suggests itself. Drawing on the well known model of Hoyle, a newly created particle of matter is followed as it is gradually accelerated by the cosmological expansion of the universe. I am suggesting that at some point in this acceleration the particle undergoes a transition in time sense and that simultaneously the other two parameters of the CPT theorem of particle physics, that is, parity and charge, will also be conjugated. For example, an electron with time sense +1and parity even will become a positron with time sense 1 and parity odd. According to Stannard's treatment, this particle, having reversed its time sense, will no longer interact with particles of a positive time sense and therefore will no longer be detectable in the part of the universe consisting of matter with a positive time sense.

Let the particle be embedded in a system which is undergoing a gravitational collapse; furthermore, let all particles in this system undergo the CPT reversal at nearly the same time, that time being coincident with the occurrence of the Schwarzschild singularity, that is, the point at which the limiting radius,  $R = GM/C^2$ , is reached. Seen from outside the collapsing system, there will be no effect, but the mass inside the singularity, rather than being lost or excluded from the universe as was necessary in previous treatments of collapse, is merely shifted from one phase of the universe to another. With the occurrence of time reversal the gravitationally contained mass reverses from a contraction to a violent expansion; this expansion takes place in a Minkowski space which is the conjugate of that in which the contraction occurs,

CPT reversal inside a gravitational singularity is believed to be a relevant issue for at least two reasons: (1) it is felt that the unique "closed-system" environment of a collapsed system is more likely to be conducive to such a reversal than any other environment; (2) it is felt that gravitational collapse is an inevitable and cosmo-